

Mathematical models and numerics for free-boundary MHD and two-phase MHD

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Mathematical model

Free-boundary MHD = MHD in one bulk + moving interface.

we consider following incompressible inviscid MHD equations in a moving domain \mathcal{D} and surface tension on the boundary

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} - \mathbf{B} \cdot \nabla \mathbf{B} + \nabla P = \mathbf{0}, \quad (1a)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (1b)$$

$$\partial_t \mathbf{B} + \mathbf{u} \cdot \nabla \mathbf{B} - \mathbf{B} \cdot \nabla \mathbf{u} = \mathbf{0}, \quad (1c)$$

$$\nabla \cdot \mathbf{B} = 0, \quad (1d)$$

where $\mathcal{D} = \cup_{t \in (0, T]} \{t\} \times \Omega_+(t)$, $\Omega_+(t) \subset \mathbb{R}^d$, and $P = p + \frac{1}{2} |\mathbf{B}|^2$ is the total pressure.

Mathematical model

We require following boundary conditions on the free boundary $\partial\mathcal{D} = \cup_{t \in (0, T]} \{t\} \times \partial\Omega_+(t)$. We will denote $\Gamma(t) = \partial\Omega_+(t)$.

- The domain is material, i.e., $(\partial_t + \mathbf{u} \cdot \nabla)|_{\partial\mathcal{D}} \in \mathcal{T}(\partial\mathcal{D})$.
Equivalently, the normal velocity of the boundary is determined by the normal velocity of fluid

$$\mathcal{V}_n = \mathbf{u} \cdot \mathbf{n}, \quad \text{on } \Gamma(t). \quad (2)$$

- Surface tension condition, i.e.,

$$P = \sigma H, \quad \text{on } \Gamma(t),$$

where σ is a positive coefficient of surface tension and H is the mean curvature of $\Gamma(t)$.

Mathematical model

- The domain is magnetic, i.e. $\mathbf{B} \cdot \mathbf{n} = 0$ on $\Gamma(t)$, where \mathbf{n} is the exterior unit normal on $\Gamma(t)$.

We further set the initial condition $\Omega_+(0) \subset \mathbb{R}^3$ is connected domain and function \mathbf{u}^0 and \mathbf{B}^0 satisfies

$$\nabla \cdot \mathbf{u}^0 = 0, \quad \nabla \cdot \mathbf{B}^0 = 0, \quad (3)$$

and $\mathbf{B}^0 \cdot \mathbf{n}|_{\Gamma(0)} = 0$.

Physical background

The free-boundary problem originates from the plasma-vacuum free-interface model, which is an important theoretic model both in the laboratory and in astrophysical magnetohydrodynamics.

The basic setting is that the plasma is confined in a vacuum $\Omega = \Omega_+(t) \cup \Omega_-(t)$ with another magnetic field $\hat{\mathbf{B}}$ and

- On $\Omega_+(t)$, the plasma is moving by the free-boundary MHD and therefore its boundary $\Gamma(t) = \partial\Omega_+$ is also moving.
- On $\Omega_-(t)$, the pre-Maxwell system holds in vacuum

$$\nabla \times \hat{\mathbf{B}} = 0, \quad \nabla \cdot \hat{\mathbf{B}} = 0. \quad (4)$$

- On the interface $\Gamma(t)$, the perfect conducting condition is required that there is no jump in the normal-component:

$$\mathbf{B} \cdot \mathbf{n} = \hat{\mathbf{B}} \cdot \mathbf{n} = 0, \quad \llbracket P \rrbracket = \sigma H. \quad (5)$$

Physical background

- There is a rigid wall wrapping the vacuum region on which the following boundary condition holds on Ω

$$\widehat{\mathbf{B}} \cdot \mathbf{n}_{\partial\Omega} = \mathbf{J}, \quad (6)$$

where \mathbf{J} is the given outer surface current density and $\mathbf{n}_{\partial\Omega}$ is the unit outer normal on $\partial\Omega$.

We remark that this setting can be viewed as one-side version of two-phase MHD.

Well-posedness, Ill-posedness and energy-conservation

The well-posedness theory of free-boundary MHD is recently investigated by Gu-Luo-Zhang and Luo-Zhang, follows from the seminal and parallel work of Coutand-Shkoller for Euler free-boundary problem. The main results are listed as follows:

- For zero surface tension case (i.e. $\sigma = 0$). A sufficient condition for the well-posedness is Rayleigh-Taylor's sign condition

$$\nabla P \cdot \mathbf{n} < 0, \quad \text{on } \Gamma(t). \quad (7)$$

This is proved by Hao and Luo (Arch. Ration. Mech. Anal., 2014).

Well-posedness, Ill-posedness and energy-conservation

- For zero surface tension case (i.e. $\sigma = 0$). If the Rayleigh-Taylor's sign condition is violated, then the problem is ill-posed, showed by Hao and Luo (Commun. Math. Phys. 2020). This is called the Rayleigh-Taylor's instability.
- The positive surface tension can stabilise the problem. The following theorem is proved by Gu, Luo and Zhang (J. Math. Pures Appl. 2024).

Well-posedness, Ill-posedness and energy-conservation

Theorem

Under suitable regularity assumption, and set $\Omega = \mathbb{T}^2 \times (0, 1)$. There exists $T > 0$ such that the above free boundary system has a unique strong solution with the energy estimate

$$\sup_{0 \leq t \leq T} \mathcal{E}(t) \leq C. \quad (8)$$

Here the model case $\Omega = \mathbb{T}^2 \times (0, 1)$ is used for their Lagrangian description in local coordinates. More explicitly, $\partial\Omega = \Gamma_0 \cup \Gamma$, where $\Gamma_0 = \mathbb{T}^2 \times \{0\}$ is a fixed bottom and $\Gamma = \mathbb{T}^2 \times \{1\}$ is the top (moving) boundary we concern.

We remark that these works follow/extend the well-posedness and ill-posedness theory of Euler free boundary problem

Well-posedness, Ill-posedness and energy-conservation

Theorem (Energy conservation)

Define the total energy

$$\mathcal{E}(t) = \frac{1}{2} \int_{\Omega_+(t)} |\mathbf{u}|^2 + |\mathbf{B}|^2 \, d\mathcal{L}^d + \sigma \int_{\Gamma(t)} 1 \, d\mathcal{H}^{d-1}. \quad (9)$$

Then it holds

$$\frac{d}{dt} \mathcal{E}(t) = 0. \quad (10)$$

Well-posedness, Ill-posedness and energy-conservation

Proof. We compute by Reynolds transport theorem and integration by parts

$$\begin{aligned} & \frac{d}{dt} \frac{1}{2} \int_{\Omega(t)} |\mathbf{u}|^2 + |\mathbf{B}|^2 \, d\mathcal{L}^d \\ &= \int_{\Omega_+(t)} (\partial_t + \mathbf{u} \cdot \nabla) \mathbf{u} \cdot \mathbf{u} + (\partial_t + \mathbf{u} \cdot \nabla) \mathbf{B} \cdot \mathbf{B} \, d\mathcal{L}^d \\ &= \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{B} \cdot \mathbf{u} - \nabla P \cdot \mathbf{u} + \mathbf{B} \cdot \nabla \mathbf{u} \cdot \mathbf{B} \, d\mathcal{L}^d \\ &= - \int_{\Gamma(t)} (\mathbf{u} \cdot \mathbf{n}) P \, d\mathcal{H}^{d-1}, \end{aligned}$$

Well-posedness, Ill-posedness and energy-conservation

where we use the divergence free property of \mathbf{u} and \mathbf{B} and its magnetic condition boundary to notice

$$\begin{aligned}\int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{u} \cdot \mathbf{B} \, d\mathcal{L}^d &= \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla (\mathbf{u} \cdot \mathbf{B}) - \mathbf{B} \cdot \nabla \mathbf{B} \cdot \mathbf{u} \, d\mathcal{L}^d \\ &= \int_{\Omega_+(t)} -\mathbf{B} \cdot \nabla \mathbf{B} \cdot \mathbf{u} \, d\mathcal{L}^d.\end{aligned}$$

Well-posedness, Ill-posedness and energy-conservation

Then we compute the surface energy is also computed by Reynolds transport theorem as

$$\frac{d}{dt} \sigma \int_{\Gamma(t)} 1 \, d\mathcal{H}^{d-1} = \sigma \int_{\Gamma(t)} \nu_{\mathbf{n}} H \, d\mathcal{H}^{d-1}.$$

Thus combine with the surface tension equation on the boundary and the property of material domain, we complete the proof.

Weak formulation

We are interested in designing a numerical method that simulate the evolution of free-boundary model of MHD. One of the key questions is:

How to represent the moving interface?

One classical way is the parametric finite element method, that is, we assume there exists a flow map from the reference initial boundary $\widehat{\Gamma}$ to the current boundary

$$\mathbf{X} : \widehat{\Gamma} \times (0, T] \rightarrow \mathbb{R}^d, \quad (11)$$

and the current boundary is given by $\Gamma(t) = \mathbf{X}(\widehat{\Gamma}, t)$ with initial condition $\Gamma(0) = \widehat{\Gamma}$.

Weak formulation

To derive a weak formulation for $(\mathbf{u}, \mathbf{B}, P)$ and \mathbf{X} , we introduce the function space

$$\mathbb{U} := [H^1(\Omega_+)]^d,$$
$$\mathbb{P} := L^2(\Omega_+), \quad \widehat{\mathbb{P}} := \left\{ q \in \mathbb{P} : \int_{\Omega_+} q \, d\mathcal{L}^d = 0 \right\}.$$

We observe that for time-dependent test function supported in $\Omega_+(t)$, we have

$$\begin{aligned} \frac{d}{dt} \int_{\Omega_+(t)} \mathbf{u} \cdot \mathbf{v} \, d\mathcal{L}^d &= \int_{\Omega_+(t)} (\partial_t + \mathbf{u} \cdot \nabla)(\mathbf{u} \cdot \mathbf{v}) \, d\mathcal{L}^d \\ &= \int_{\Omega_+(t)} \partial_t \mathbf{u} \cdot \mathbf{v} + \mathbf{u} \cdot \partial_t \mathbf{v} - (\nabla \cdot \mathbf{u})(\mathbf{u} \cdot \mathbf{v}) \, d\mathcal{L}^d \\ &\quad + \int_{\Gamma(t)} (\mathbf{u} \cdot \mathbf{n})(\mathbf{u} \cdot \mathbf{v}) \, d\mathcal{H}^{d-1}, \end{aligned}$$

Weak formulation

thus combine with the divergence-free condition of \mathbf{u} , we can derive

$$\begin{aligned} & \int_{\Omega_+(t)} \partial_t \mathbf{u} \cdot \mathbf{v} \, d\mathcal{L}^d \\ &= \frac{1}{2} \left[\frac{d}{dt} \int_{\Omega_+(t)} \mathbf{u} \cdot \mathbf{v} \, d\mathcal{L}^d + \int_{\Omega_+(t)} \partial_t \mathbf{u} \cdot \mathbf{v} - \mathbf{u} \cdot \partial_t \mathbf{v} \, d\mathcal{L}^d \right] \\ & \quad - \frac{1}{2} \int_{\Gamma(t)} (\mathbf{u} \cdot \mathbf{n})(\mathbf{u} \cdot \mathbf{v}) \, d\mathcal{H}^{d-1} \\ &=: \frac{1}{2} \frac{d}{dt} (\mathbf{u}, \mathbf{v})_{\Omega_+} + \frac{1}{2} d_{\Omega_+} (\mathbf{u}, \mathbf{v}) - \frac{1}{2} \langle \mathbf{u} \cdot \mathbf{n}, \mathbf{u} \cdot \mathbf{v} \rangle_{\Gamma}. \end{aligned}$$

Weak formulation

Then, by using the general identity for any vector valued function $\mathbf{u}, \mathbf{v}, \mathbf{w}$

$$(\mathbf{u} \cdot \nabla) \mathbf{v} \cdot \mathbf{w} = \frac{1}{2} [(\mathbf{u} \cdot \nabla) \mathbf{v} \cdot \mathbf{w} - (\mathbf{u} \cdot \nabla) \mathbf{w} \cdot \mathbf{v}] + \frac{1}{2} (\mathbf{u} \cdot \nabla) (\mathbf{v} \cdot \mathbf{w}),$$

we obtain

$$\begin{aligned} \int_{\Omega_+(t)} (\mathbf{u} \cdot \nabla) \mathbf{u} \cdot \mathbf{v} \, d\mathcal{L}^d &= \frac{1}{2} \left[\int_{\Omega_+(t)} (\mathbf{u} \cdot \nabla) \mathbf{u} \cdot \mathbf{v} - (\mathbf{u} \cdot \nabla) \mathbf{v} \cdot \mathbf{u} \, d\mathcal{L}^d \right] \\ &\quad + \frac{1}{2} \int_{\Gamma(t)} (\mathbf{u} \cdot \mathbf{n}) (\mathbf{u} \cdot \mathbf{v}) \, d\mathcal{H}^{d-1} \\ &=: \frac{1}{2} c_{\Omega_+}(\mathbf{u}, \mathbf{u}, \mathbf{v}) + \frac{1}{2} \langle \mathbf{u} \cdot \mathbf{n}, \mathbf{u} \cdot \mathbf{v} \rangle_{\Gamma} \end{aligned}$$

Weak formulation

Thus, the material derivative term can be written as

$$\begin{aligned} & \int_{\Omega_+(t)} (\partial_t + \mathbf{u} \cdot \nabla) \mathbf{u} \cdot \mathbf{v} \, d\mathcal{L}^d \\ &= \frac{1}{2} \frac{d}{dt} (\mathbf{u}, \mathbf{v})_{\Omega_+} + \frac{1}{2} d_{\Omega_+} (\mathbf{u}, \mathbf{v}) + \frac{1}{2} c_{\Omega_+} (\mathbf{u}, \mathbf{u}, \mathbf{v}). \end{aligned}$$

Similarly, for the magnetic field term can be written as

$$\begin{aligned} & \int_{\Omega_+(t)} (\partial_t + \mathbf{u} \cdot \nabla) \mathbf{B} \cdot \mathbf{C} \, d\mathcal{L}^d \\ &= \frac{1}{2} \frac{d}{dt} (\mathbf{B}, \mathbf{C})_{\Omega_+} + \frac{1}{2} d_{\Omega_+} (\mathbf{B}, \mathbf{C}) + \frac{1}{2} c_{\Omega_+} (\mathbf{B}, \mathbf{B}, \mathbf{C}). \end{aligned}$$

Weak formulation

Subsequently, we will deal with the additional terms. Indeed, by using magnetic boundary condition, we write

$$\begin{aligned}
 & \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{B} \cdot \mathbf{v} \, d\mathcal{L}^d \\
 &= \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla (\mathbf{B} \cdot \mathbf{v}) \, d\mathcal{L}^d - \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{v} \cdot \mathbf{B} \, d\mathcal{L}^d \\
 &= - \int_{\Omega_+(t)} (\nabla \cdot \mathbf{B})(\mathbf{B} \cdot \mathbf{v}) \, d\mathcal{L}^d + \int_{\Gamma(t)} (\mathbf{B} \cdot \mathbf{n})(\mathbf{B} \cdot \mathbf{v}) \, d\mathcal{H}^{d-1} \\
 &\quad - \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{v} \cdot \mathbf{B} \, d\mathcal{L}^d \\
 &= - \int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{v} \cdot \mathbf{B} \, d\mathcal{L}^d \\
 &=: -\ell_{\Omega_+(t)}(\mathbf{B}, \mathbf{v}, \mathbf{B}),
 \end{aligned}$$

Weak formulation

and without any modification, we have

$$\int_{\Omega_+(t)} \mathbf{B} \cdot \nabla \mathbf{u} \cdot \mathbf{C} \, d\mathcal{L}^d = \ell_{\Omega_+(t)}(\mathbf{B}, \mathbf{u}, \mathbf{C}).$$

Finally for the treatment of total pressure term, we pair the term ∇P with the test \mathbf{u} and integration by parts and together with the surface tension condition

$$\begin{aligned} \int_{\Omega_+(t)} \nabla P \cdot \mathbf{v} \, d\mathcal{L}^d &= - \int_{\Omega_+(t)} (\nabla \cdot \mathbf{v}) P \, d\mathcal{L}^d + \int_{\Gamma(t)} (\mathbf{v} \cdot \mathbf{n}) P \, d\mathcal{H}^{d-1} \\ &=: -(\nabla \cdot \mathbf{v}, P)_{\Omega_+} + \sigma \langle \mathbf{v} \cdot \mathbf{n}, H \rangle_{\Gamma}. \end{aligned}$$

Weak formulation

The representation or computation of the mean curvature is further depending on how we represent the interface. In the parametric setting, we have a classical formula for mean curvature

$$H\mathbf{n} = \Delta_{\Gamma}\text{Id}, \quad (12)$$

where Δ_{Γ} is the surface Laplacian and Id is viewed as an identity function on Γ . To consider a weak formulation of surface, we use the surface function spaces

$$\mathbb{V} = H^1(\Gamma), \quad \mathbb{W} = L^2(\Gamma)$$

Weak formulation

and use the assumption that the normal velocity of interface is controlled by the fluid, we get for any $(\varphi, \boldsymbol{\omega}) \in \mathbb{W} \times \mathbb{V}$

$$\begin{aligned}\langle \partial_t \mathbf{X} \cdot \mathbf{n}, \varphi \rangle_\Gamma - \langle \mathbf{u} \cdot \mathbf{n}, \varphi \rangle_\Gamma &= 0, \\ \langle H\mathbf{n}, \boldsymbol{\omega} \rangle_\Gamma + \langle \nabla_\Gamma \text{Id}, \nabla_\Gamma \boldsymbol{\omega} \rangle_\Gamma &= 0.\end{aligned}$$

Weak formulation

Sum above derivations up, we finally get the full weak formulation:
Find $(\mathbf{u}(\cdot, t), \mathbf{B}(\cdot, t), P(\cdot, t)) \in \mathbb{U} \times \mathbb{U} \times \widehat{\mathbb{P}}$, and $\mathbf{X}(\cdot, t) \in [\mathbb{V}]^3$ and $H(\cdot, t) \in \mathbb{W}$ such that

$$0 = \frac{1}{2} \frac{d}{dt} (\mathbf{u}, \mathbf{v})_{\Omega_+} + \frac{1}{2} d_{\Omega_+} (\mathbf{u}, \mathbf{v}) + \frac{1}{2} c_{\Omega_+} (\mathbf{u}, \mathbf{u}, \mathbf{v}) + \ell_{\Omega_+(t)} (\mathbf{B}, \mathbf{v}, \mathbf{B}) \quad (13a)$$

$$- (\nabla \cdot \mathbf{v}, P)_{\Omega_+} + \sigma \langle \mathbf{u} \cdot \mathbf{n}, H \rangle_{\Gamma}, \quad (13b)$$

$$0 = (\nabla \cdot \mathbf{u}, q)_{\Omega_+} \quad (13c)$$

$$0 = \frac{1}{2} \frac{d}{dt} (\mathbf{B}, \mathbf{C})_{\Omega_+} + \frac{1}{2} d_{\Omega_+} (\mathbf{B}, \mathbf{C}) + \frac{1}{2} c_{\Omega_+} (\mathbf{B}, \mathbf{B}, \mathbf{C}) - \ell_{\Omega_+(t)} (\mathbf{B}, \mathbf{u}, \mathbf{C}) \quad (13d)$$

$$0 = \langle \partial_t \mathbf{X} \cdot \mathbf{n}, \varphi \rangle_{\Gamma} - \langle \mathbf{u} \cdot \mathbf{n}, \varphi \rangle_{\Gamma}, \quad (13e)$$

$$0 = \langle H \mathbf{n}, \boldsymbol{\omega} \rangle_{\Gamma} + \langle \nabla_{\Gamma} \text{Id}, \nabla_{\Gamma} \boldsymbol{\omega} \rangle_{\Gamma}, \quad (13f)$$

for any $(\mathbf{v}, \mathbf{C}, q) \in \mathbb{U} \times \mathbb{U} \times \widehat{\mathbb{P}}$ and $(\varphi, \boldsymbol{\omega}) \in \mathbb{W} \times [\mathbb{V}]^3$.

Weak formulation

Theorem

Let $(\mathbf{u}(\cdot, t), \mathbf{B}(\cdot, t), P(\cdot, t))$, and $(\mathbf{X}(\cdot, t), H(\cdot, t))$ is a solution of above system. Then it holds

$$\frac{d}{dt} \mathcal{E}(t) = 0. \quad (14)$$

Here the energy is defined by

$$\mathcal{E}(t) = \frac{1}{2} \int_{\Omega_+} |\mathbf{u}|^2 + |\mathbf{B}|^2 \, d\mathcal{L}^d + \sigma \int_{\Gamma(t)} 1 \, d\mathcal{H}^{d-1}. \quad (15)$$

The proof is essentially the same in the semi-discrete case, so we postpone the proof later.

Finite element approximation

We now consider the finite element discretization of above system.
For the moving interface we can use

- The unfitted mesh approach. i.e., the interface and the bulk meshes are completely independent of each other.
- The fitted mesh approach, it ensures that the bulk mesh is aligned with the interface.

Finite element approximation

We start with the semi-discretization. Assume that $\Omega^h \subset \mathbb{R}^d$, $d \in \{2, 3\}$, is a bounded Lipschitz domain with polygonal (or polyhedral, respectively) boundary, and $\Gamma^h \subset \Omega^h$ is a $(d - 1)$ -dimensional polyhedral surface, fully contained in the interior of Ω . We suppose there exist decompositions

$$\overline{\Omega^h} = \bigcup_{j=1}^{J_{\Omega^h}} \overline{\sigma_j}, \quad \Gamma^h = \bigcup_{j=1}^{J_{\Gamma^h}} \overline{\sigma_j},$$

where $\{\sigma_j\}_{j=1}^{J_{\Omega^h}}$ and $\{\sigma_j\}_{j=1}^{J_{\Gamma^h}}$ are families of mutually disjoint open d -simplices and $(d - 1)$ -simplices, respectively.

Finite element approximation

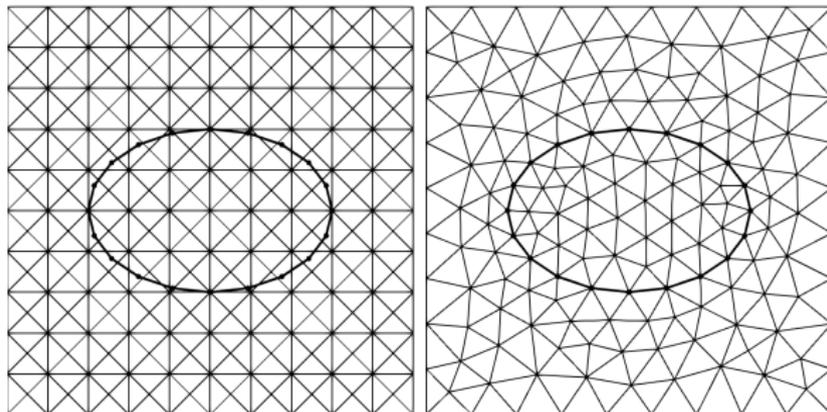


Figure: Unfitted and fitted bulk meshes together with an interface mesh.

Finite element approximation

We define the finite element spaces on Ω^h by

$$S^k(\mathcal{T}^h) := \{\chi \in C(\overline{\Omega}^h) : \chi|_o \in P_k(o), \forall o \in \mathcal{T}^h\} \subset W^{1,\infty}(\Omega^h), \quad k \geq 1,$$

$$S^0(\mathcal{T}^h) := \{\chi \in L^\infty(\Omega^h) : \chi|_o \in P_0(o), \forall o \in \mathcal{T}^h\} \subset L^\infty(\Omega^h),$$

where $P^k(o)$ denotes the space of polynomials of degree less than or equal to $k \in \mathbb{N}_{\geq 0}$ on $o \in \mathcal{T}^h$.

Finite element approximation

Given Ω^h and Γ^h , we can separate Ω^h into the interior part exterior part

$$\Omega^h = \overline{\Omega_+^h} \cup \overline{\Omega_-^h}, \quad (16)$$

and defined the restricted finite element space on $\overline{\Omega_+^h}$

$$S_+^k(\mathcal{T}^h) = \left\{ \chi \in S^k(\mathcal{T}^h) : \chi(\mathbf{p}_i) = 0 \text{ if } \text{supp}(\chi) \subset \overline{\Omega_-^h}, i = 1, \dots, K_{S^k} \right\}. \quad (17)$$

The inner product on $(\cdot, \cdot)_{\Omega_+^h}$ is naturally induced.

Finite element approximation

We choose the discrete function space tuple $(\mathbf{U}^h(\mathcal{T}^h), \hat{\mathbb{P}}^h(\mathcal{T}^h)) \subset (\mathbf{U}, \hat{\mathbb{P}})$ for velocity/pressure such that the usual inf-sup condition is satisfied, meaning there exists a constant $C_0 > 0$ independent of h such that

$$\inf_{\varphi \in \hat{\mathbb{P}}^h(\mathcal{T}^h)} \sup_{\boldsymbol{\xi} \in \mathbf{U}^h(\mathcal{T}^h)} \frac{(\varphi, \nabla \cdot \boldsymbol{\xi})}{\|\varphi\|_{L^2} \|\boldsymbol{\xi}\|_{H^1}} \geq C_0 > 0. \quad (18)$$

For instance, we can choose the lowest order Taylor-Hood element.

Finite element approximation

The finite element space on Γ^h is defined by

$$\mathbb{V}^h(\Gamma^h) \{ \chi \in C(\Gamma^h) : \chi|_{\sigma_j} \in P^1(\sigma_j), j = 1, \dots, J_{\Gamma^h} \} \subset H^1(\Gamma^h),$$

where $P^1(\sigma_j)$ denotes the space of polynomials of degree 1 on σ_j , $j = 1, \dots, J_{\Gamma^h}$. We define the L^2 -inner product and its lumped mass version on polyhedral surface Γ^h as follows

$$\langle \mathbf{u}, \mathbf{v} \rangle_{\Gamma^h} := \int_{\Gamma^h} \mathbf{u} \cdot \mathbf{v} \, d\mathcal{H}^{d-1}.$$

Finite element approximation

Based on above construction, the semi-discrete scheme can be described as follows: Given initial approximation $\Gamma^h(0)$ of $\Gamma(0)$ and initial velocity $\mathbf{u}^h(0)$, find

$$(\mathbf{u}^h(\cdot, t), \mathbf{B}^h(\cdot, t), P^h(\cdot, t)) \in \mathbb{U}^h(\mathcal{T}^h) \times \mathbb{U}^h(\mathcal{T}^h) \times \widehat{\mathbb{P}}(\mathcal{T}^h),$$

and

$$\mathbf{X}^h(\cdot, t) \in [\mathbb{V}^h(\Gamma^h)]^3, \quad \text{and} \quad H^h(\cdot, t) \in \mathbb{V}^h(\Gamma^h),$$

such that for any $(\mathbf{v}^h, \mathbf{C}^h, q^h) \in \mathbb{U}^h(\mathcal{T}^h) \times \mathbb{U}^h(\mathcal{T}^h) \times \widehat{\mathbb{P}}(\mathcal{T}^h)$ and $(\varphi^h, \boldsymbol{\omega}^h) \in \mathbb{V}^h(\Gamma^h) \times [\mathbb{V}^h(\Gamma^h)]^3$.

Finite element approximation

$$0 = \frac{1}{2} \frac{d}{dt} (\mathbf{u}^h, \mathbf{v}^h)_{\Omega_+^h} + \frac{1}{2} d_{\Omega_+^h} (\mathbf{u}^h, \mathbf{v}^h) \quad (19a)$$

$$+ \frac{1}{2} c_{\Omega_+^h} (\mathbf{u}^h, \mathbf{u}^h, \mathbf{v}^h) + \ell_{\Omega_+^h} (\mathbf{B}^h, \mathbf{v}^h, \mathbf{B}^h) - (\nabla \cdot \mathbf{v}^h, P^h)_{\Omega_+^h} + \sigma \langle \mathbf{u}^h \cdot \mathbf{n}^h, H^h \rangle_{\Gamma^h}, \quad (19b)$$

$$0 = (\nabla \cdot \mathbf{u}^h, q^h)_{\Omega_+^h} \quad (19c)$$

$$0 = \frac{1}{2} \frac{d}{dt} (\mathbf{B}^h, \mathbf{C}^h)_{\Omega_+^h} + \frac{1}{2} d_{\Omega_+^h} (\mathbf{B}^h, \mathbf{C}^h) + \frac{1}{2} c_{\Omega_+^h} (\mathbf{B}^h, \mathbf{B}^h, \mathbf{C}^h) - \ell_{\Omega_+^h(t)} (\mathbf{B}^h, \mathbf{u}^h, \mathbf{C}^h) \quad (19d)$$

$$0 = \langle \partial_t \mathbf{X}^h \cdot \mathbf{n}^h, \varphi^h \rangle_{\Gamma^h} - \langle \mathbf{u}^h \cdot \mathbf{n}^h, \varphi^h \rangle_{\Gamma^h}, \quad (19e)$$

$$0 = \langle H^h \mathbf{n}^h, \omega^h \rangle_{\Gamma^h} + \langle \nabla_{\Gamma^h} \text{Id}, \nabla_{\Gamma^h} \omega^h \rangle_{\Gamma^h}, \quad (19f)$$

Finite element approximation

Theorem

Let $(\mathbf{u}^h(\cdot, t), \mathbf{B}^h(\cdot, t), P^h(\cdot, t))$, and $(\mathbf{X}^h(\cdot, t), H^h(\cdot, t))$ is a solution of above system. Then it holds

$$\frac{d}{dt} \mathcal{E}^h(t) = 0. \quad (20)$$

Here the energy is defined by

$$\mathcal{E}^h(t) = \frac{1}{2} \int_{\Omega_+^h} |\mathbf{u}^h|^2 + |\mathbf{B}^h|^2 \, d\mathcal{L}^d + \sigma \int_{\Gamma^h(t)} 1 \, d\mathcal{H}^{d-1}. \quad (21)$$

Finite element approximation

Proof. Take test function $\mathbf{v}^h = \mathbf{u}^h$, $\mathbf{C}^h = \mathbf{B}^h$, we get

$$\begin{aligned} & \frac{d}{dt} \frac{1}{2} \left[\int_{\Omega_+^h} |\mathbf{u}^h|^2 + |\mathbf{B}^h|^2 \, d\mathcal{L}^d \right] \\ &= \frac{1}{2} \frac{d}{dt} (\mathbf{u}^h, \mathbf{u}^h)_{\Omega_+^h} + \frac{1}{2} \frac{d}{dt} (\mathbf{B}^h, \mathbf{B}^h)_{\Omega_+^h} \\ &= -\frac{1}{2} d_{\Omega_+^h} (\mathbf{u}^h, \mathbf{u}^h) - \frac{1}{2} c_{\Omega_+^h} (\mathbf{u}^h, \mathbf{u}^h, \mathbf{u}^h) - \ell_{\Omega_+^h} (\mathbf{B}^h, \mathbf{u}^h, \mathbf{B}^h) + (\nabla \cdot \mathbf{u}^h, P^h)_{\Omega_+^h} \\ & \quad - \sigma \langle \mathbf{u}^h \cdot \mathbf{n}^h, H^h \rangle_{\Gamma^h} - \frac{1}{2} d_{\Omega_+^h} (\mathbf{B}^h, \mathbf{B}^h) - \frac{1}{2} c_{\Omega_+^h} (\mathbf{B}^h, \mathbf{B}^h, \mathbf{B}^h) + \ell_{\Omega_+^h(t)} (\mathbf{B}^h, \mathbf{u}^h, \mathbf{B}^h), \end{aligned}$$

Finite element approximation

then combine with the fact that $d_{\Omega_+^h}(\cdot, \cdot)$ is anti-symmetric, $c_{\Omega_+^h}(\cdot, \cdot, \cdot)$ is anti-symmetric with last two components, and take the test function $q^h = P^h$ in the second equation, we arrived at

$$\begin{aligned} & \frac{d}{dt} \frac{1}{2} \left[\int_{\Omega_+^h} |\mathbf{u}^h|^2 + |\mathbf{B}^h|^2 \, d\mathcal{L}^d \right] \\ &= -\sigma \left\langle \mathbf{u}^h \cdot \mathbf{n}^h, H^h \right\rangle_{\Gamma^h}. \end{aligned}$$

Finite element approximation

Subsequently, we will use two interface equations by computing the discrete surface energy by Reynolds transport theorem and taking test function $\boldsymbol{\omega}^h = \partial_t \mathbf{X}^h$ and $\varphi^h = H^h$.

$$\begin{aligned} \frac{d}{dt} \int_{\Gamma^h(t)} 1 \, d\mathcal{H}^{d-1} &= \int_{\Gamma^h(t)} \nabla_{\Gamma^h} \text{Id} \cdot \nabla_{\Gamma^h} \partial_t \mathbf{X}^h \, d\mathcal{H}^{d-1} \\ &= \left\langle \nabla_{\Gamma^h} \text{Id}, \nabla_{\Gamma^h} \partial_t \mathbf{X}^h \right\rangle_{\Gamma^h} \\ &= - \left\langle H^h \mathbf{n}^h, \partial_t \mathbf{X}^h \right\rangle_{\Gamma^h} \\ &= \left\langle \mathbf{u}^h \cdot \mathbf{n}^h, H^h \right\rangle_{\Gamma^h}. \end{aligned}$$

To summarise, we obtain

$$\frac{d}{dt} \mathcal{E}^h = \frac{d}{dt} \frac{1}{2} \left[\int_{\Omega_+^h} |\mathbf{u}^h|^2 + |\mathbf{B}^h|^2 \, d\mathcal{L}^d \right] + \sigma \frac{d}{dt} \int_{\Gamma^h(t)} 1 \, d\mathcal{H}^{d-1} = 0.$$

Summary

Questions:

- What about full discrete case?
- How about use finite element exterior calculus such as H_{div} and H_{curl} space?
- How about the two-phase MHD model?

Thanks for Listening! Any Questions?